Data Intensive Radio Astronomy en route to the SKA

The rise of **BIG RADIO DATA**

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Technology Challenges en route to SKA

A SUUARE KILOMETRE ARRA

- Large collecting Area
- Large field-of-view
- Large bandwidth
- Power consumption
- Massive Data





- Until recently bandwidths of a few 10's of MHz were standard
 - VLA: 50 MHz per IF
- New telescopes and upgrades dramatically increase bandwidth
 - Arecibo: 300 MHz
 - ASKAP: 300 MHz
 - MeerKAT: 770 MHz
 - JVLA: 8 GHz
- At the same time observing in "spectral-line" mode
 - for RFI detection and excision (BW extends beyond protected regions)
 - Avoid signal degradation for wide-field imaging.
 - Use of multiple frequencies for visibility sampling (multi-frequency synthesis)





Jansky VLA-VLA Comparison



Parameter	VLA	JVLA	Factor	Current
Point Source Cont. Sensitivity (1 σ ,12hr.)	10 μJy	1 μJy	10	2 μJy
Maximum BW in each polarization	0.1 GHz	8 GHz	80	2 GHz
# of frequency channels at max. BW	16	16,384	1024	4096
Maximum number of freq. channels	512	4,194,304	8192	12,288
Coarsest frequency resolution	50 MHz	2 MHz	25	2 MHz
Finest frequency resolution	381 Hz	0.12 Hz	3180	.12 Hz
# of full-polarization spectral windows	2	64	32	16
(Log) Frequency Coverage (1 – 50 GHz)	22%	100%	5	100%

Field of View technologies



- Large N small D
 - 15 m antennas 1 sq deg at 1.4 GHz.



- Arecibo & FAST
 - Multiple horn receivers





- Aperture plane arrays
 - Multiple independent beams over a hemisphere



- Phased Array Feeds
 - Multiple beams providing large FOV







Sociology of Radio Astronomy



 Much of the key science en route to the SKA will be achieved via large-scale survey mode observing programs executed by globally distributed teams of researchers



Arecibo ALFA Surveys



- ALFALFA (HI)
- PALFA (Pulsars)
- GALFACTS (Spectro-polarimetry)

GALFACTS and PALFA Aggregate rate 500 MB/s



LOFAR Survey Science

- Sky surveys at 15, 30, 60, 120, 200 MHz
 - Galaxy formation
 - Intergalactic magnetic fields
 - Star formation in early universe
 - Expansion of discovery parameter space





ASKAP Survey Science

- WALLABY (HI emission)
- EMU (continuum)
- POSSUM (polarization)
- FLASH (HI absorption)
- VAST (slow transients and variables)
- GASKAP (Galactic HI)
- CRAFT (fast transients)
- DINGO (Deep HI)
- COAST (pulsar and timing survey)
- VLBI (high resolution science)

Some project will be commensal





MeerKAT Survey Science

- Pulsar Timing
- LADUMA (Deep HI)
- MESMER (High-z CO)
- MeerKAT Absorption Line Survey
- MHONGOOSE (Nearby HI)
- TRAPUM (pulsar search)
- MeerKAT HI Survey of Fornax
- MeerGAL (Galactic Plane Survey)
- MIGHTEE (Deep continuum and polarization)
- ThunderKAT (variables and transients)

Some project will be commensal





MeerKAT Large Surveys (43,000 hours allocated)



22 countries

SKA Timeline







Data, data, data, informa.....(gulp).



Survey Raw Data Rates



SQUARE KILOMETRE ARRA

The Challenge



- Survey mode observations drive:
 - Very high data rates and volumes
 - Storage, transfer, access
 - Delivery of data to end users not practical
 - Complex, multi-purpose, processing and analysis
 - Processing, analysis, visualization, data mining
 - Multiple processing and analysis chains
 - Collaborative execution by globally distributed teams of researchers
 - Distributed and remote science community
 - Distributed collaboration in data processing, analysis and science



- Dedicated observatory pipelines to
 - create science ready output at the observatory and send finished products to users
 - Being adopted by LOFAR, EVLA, ASKAP,...
- Cons
 - Monolothic system under control of central "authority" of experts
 - One-pass through data. No room for iteration and improvement
 - Disconnects the user community from intellectual development processing techniques and technologies



An alternate approach: cyberinfrastructure platform

- Use cloud and web 2.0 technologies to empower the end user
- Turn global resources into the solution
 - On-line interactive access to a global HPC cloud and data
 - Collaborative development of novel pipelines processes
 - Visualization and visual analytics of very large data
 - Creation and ingestion of derived and ancilliary data products
 - E-science tools for distributed collaboration and analysis

The CyberSKA Project



Initiative to develop a scalable and distributed cyberinfrastructure platform to meet evolving needs of data-intensive radio astronomy en route to the SKA



CyberSKA Development Partners











Focus Areas



- Collaboration
 - Portal built on social networking technologies
- Data Management
 - Scalable collaborative access, sharing and searching of distributed (BIG) data sets
- Data Processing
 - Framework for executing algorithms and workflows
- Data Visualization and Visual analytics

 On-line interactive visualization of remote Big Data
- Third Party Applications
 - Community driven site with common API



CyberSKA High Level Architecture



Collaborative Portal



- Portal built on top of the Elgg open source social networking platform
 - Provides many features including: tags, bookmarks, profiles, blogs, wikis, contacts, groups, document sharing, discussions, messaging, calendars, status, activity feeds



Distributed Data System



- Based on iRODS (Integrated Rule-Oriented Data System)
 - Abstracts data location
 - Supports data replication / cross-site backup
 - Efficient WAN data transfer
 - Rule engine to automate various tasks
- Upload/download tools
 - Java Applet / Java Web Start based
 - Supports "large" data uploads/downloads
- Automated mime type recognition
 - For many common file types
 - FITS and Measurement Set (CASA) image data or visibility data
- Automated header extraction and thumbnail generation





On-line visualization of Big Data



VM based on-line interactive visual analytics of large, multi-dimensional image cubes

- 0.5 TB full Stokes I, Q, U image cube sets
- Collaboration, screen sharing, etc.



Third Party Application Interface



- API for integrating third party / "remote" server applications
- Single sign-on to applications enabled using Oauth
- Push/pull information and data to/from portal
- Current applications include PALFA Candidate Viewer, PALFA Top Candidates, PALFA Observation Scheduler, PALFA Diagnostics Tool, GALFACTS Processing Pipeline, Visibility Data Processing Pipeline, Source Counts, …

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IVOA Data Query Interface

CyberSKA

User Data

User

Data Ingest Service

Data Management Service



CyberSKA IVOA Collection



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CyberSKA Usage



- 258 members from around the world
- 40+ "groups" (GALFACTS, PALFA, RM Synthesis, EVLA Deep Polarization Field, GMRT Deep Polarization Field, CASA Users, ...)



Imaging Survey Co-development





- GALFACTS (Arecibo) 49 collaborators on CyberSKA
- Deep Polarization Field Surveys (EVLA, GMRT)
- CyberSKA:
 - Used for sharing documents, creating wiki pages, having discussions and bookmarking resources
 - Enables on-line visualization of remote data sets
 - Access to GALFACTS survey data and third party applications for running data processing pipelines

Pulsar Survey Co-development

- PALFA (Arecibo) 69 collaborators
- CyberSKA:
 - shared hub for documentation, meeting minutes, publications and task lists
 - Used as an on-line application centre for single sign-on access to a variety of third-party applications
 - Resulted in a significant increase in the rate of discoveries







On-going and future work



- Study funded by North American ARC in collaboration with Harvard to adapt on-line visual analytics for ALMA on-line data system
- Completion of user interface for data pipeline tool to allow user developed pipeline processing
- Collaboration with CADC on next generation IVOA interface to include visibility data sets (CAOM2)
- Incorporation of distributed HPC-based cloud architecture for scalability to multi-site petascale (Compute Canada, IBM Watson).

The Square Kilometre Array

A Global Observatory

A Global Solution to BIG Data